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The emission lines in Be stars are thought to originate from a low-latitude disk or ring-like envelope (Kogure & Leung 2007), which in the case of BD+50 3188 is measured to be rotating at 138 km s...

*The Astrophysics of Emission-Line Stars - ResearchGate*

Astrophysics > Astrophysics of Galaxies. arXiv:2009.04193 (astro-ph) [Submitted on 9 Sep 2020] Title: A Broad Absorption Line Outflow Associated with the Broad Emission Line Region in the Quasar SDSS J075133.35+134548.3. Authors: Bo Liu, Hongyan Zhou, Xinwen Shu, ...

*[2009.04193] A Broad Absorption Line Outflow Associated ...*

Abstract: The next generation of spectroscopic surveys will target emission-line galaxies (ELGs) to produce constraints on cosmological parameters. We study the large scale structure traced by ELGs using a combination of a semi-analytical model of galaxy formation, a code that computes the nebular emission from HII regions using the properties of the interstellar medium, and a large-volume, high-resolution N-body simulation.

*[2010.08500] The assembly bias of emission line galaxies*

The Astrophysics of Emission Line Stars offers general information on emission line stars, starting from a brief introduction to stellar astrophysics, and then moving toward a broad overview of emission line stars including early and late type stars as well as pre-main sequence stars. Detailed references have been prepared along with an index for further reading.

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This causes line emission spectra to be produced, as shown below. This is not a continuous spectrum as only light of specific frequencies and specific colours are produced. Different types of atoms...

*Emission spectra - Spectra - Higher Physics Revision - BBC ...*

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[2010.06902] *The X-SHOOTER/ALMA sample of Quasars in the ...*

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Aug 31, 2020 the astrophysics of emission line stars astrophysics and space science library Posted By Frank G. SlaughterPublic Library TEXT ID 578acc0d Online PDF Ebook Epub Library a study on the photosphere and disc of the be star psc is presented we recover almost 40 years of high resolution spectroscopic observations and additional data gathered from the bess database we evaluate

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In general, emission lines occur in gases that are optically thin, or in atmospheres that have a temperature inversion (i.e. are hotter closer to the observer). In an optically thin gas, if excited atoms or ions are radiatively deexcited, then the photons produced can escape to an observer.

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